Stability of SD energy reconstructions under NKG and Power Law lateral distribution functions

Update since Lisbon 2013



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Background

Should surface detector energy reconstructions depend on the lateral distribution function (LDF)?

IDEALLY, NO

By design, S1000 is least sensitive to different parameterizations of the LDF. *(Newton, Knapp, Watson 2006)*

REALISTICALLY, SOME

- CIC(θ) is LDF dependent. (Schmidt, Maris, Roth 2006)
- Core reconstruction is LDF dependent.
- Functions differ in their fitting behaviors.



TAKEAWAY

Appropriate, well-fitting, LDFs should yield similar values for S1000, and in turn energy, as long as S1000 is well constrained. Small differences, however, are expected.

Previously Reported Results

DATA SET: 01.2004 – 08.2012 **OFFLINE VERSION:** 2.7.8 **PHYSICAL TRIGGER:** 6T5

LDFS: NKG and Power Law

Both functions have been substantiated and fit the distance vs. signal data well.



RESULTS:

- Significant differences in energy reconstructions at low zenith angles due to poor bracketing of \$1000.
- Dominance of such significantly different events at high energies.

PROPOSED MODIFICATIONS:

- Derive independent CIC(θ) curves for NKG and Power Law.
- Perform independent energy calibrations for NKG and Power Law.

LDFS: NKG and Power Law

Both functions have been substantiated and fit the distance vs. signal data well.

DATA SET: 2004 – 2012

OFFLINE VERSION: 2.9.1 (including ICRC2013 updates)

PHYSICAL TRIGGER: 6T5

- Separate CIC(θ) derived for NKG and Power Law (see backup slides)
- Separate Energy Calibrations performed for NKG and Power Law (see backup slides)















Bracketing Span

A UBS of 1000 m equates to a tank at 2000 m from the core, with no tanks between 1000 and 2000 m from the core.





A LBS of 1000 m equates to the shower core landing directly on the hot tank with no other tanks between the hot tank and 1000 m from the core.

Bracketing Span

If, for a shower with a low zenith angle, a tank saturates within 200 m of the core and is not recovered, the first station usable in the LDF fit is \sim 1400 m from the core.





Bracketing Span & Percent Energy Difference



$5 \text{ EeV} \le E_{Av} \le 10 \text{ EeV}$



$10 \text{ EeV} \le \text{E}_{AV} \le 20 \text{ EeV}$



$20 \text{ EeV} \le \text{E}_{Av} \le 50 \text{ EeV}$



 $E_{Av} \ge 50 \text{ EeV}$



Comparison with Observer SD Energy Uncertainties





CONCLUSIONS

- Geometry of the array coupled with shower geometries can result in poor or good bracketing of S1000.
- E_{SD} deviates significantly between reconstructions using valid, yet different LDFs
- E_{sD} differences are most prevalent for low zenith angle showers in which one or more tanks saturate.

IMPLICATIONS

- Larger uncertainty in energy.
- Possible biasing of energy calibration by events with saturated tanks.
- Possible biasing of energy spectrum, anisotropy, etc.

POSSIBLE SOLUTIONS

- Better fitting LDF which is not systematically biased for larger bracketing spans.
 - Possibility: Adelaide LDF (Alexander Herve) GAP-2013-076
- Use of S1500 for events where S1000 is typically poorly bracketed (e.g. low zenith angle, 1+ tanks saturate)

END

ADDITIONAL SLIDES



NKG





Energy Calibrations

NKG

Power Law



<u>ICRC 2013 (NKG)</u> A = 0.190 +/- 0.005 B = 1.025 +/- 0.007

Note:

- 2 FD cuts failed (minBackgroundRMS & profileChi2Sigma

- Chi squared minimization used vs. max likelihood

- Resolution in progress

 $5 \le E < 10 \text{ EeV}$ $10 \le E < 20 \text{ EeV}$ $20 \le E < 50 \text{ EeV}$ $E \ge 50 \text{ EeV}$

